COLLECTIVE EXERCISE ON IRRADIATION AND MICROSTRUCTURE EVOLUTION

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Summary of the caracteristic simulation sequence



Upon interaction with a neutron, an atom can be kicked out of its lattice position if the energy transferred is superior to the threshold displacement energy E_d



The IRRAD Module





 Norgett Robinson Torrens displacement per atom (NRT-DPA) measure the level of irradiation considering binary collision model including ballistic processes leading to recombination.[1]



[1] Norgett, M.J., M.T. Robinson and I.M. Torrens (1975), "A proposed method of calculating displacement dose rates", Nucl. Engr. and Design, 33(1):50-54.

Towards a realistic representation of primary

Damage produces during irradiation is significantly lower than the one predicted by NRT dpa.



\rightarrow Molecular Dynamics calculation



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- 200 fs: creation of large amount of damage corresponding to NRT dpa.
- Few ps: Heat spike creating a liquidlike region of high kinetic energy followed by recrystallisation → athermal recombination fewer defect
- Several s: cascade annealing perform with Kinetic Monte Carlo to account for thermal recombination of defects: cascade escape.

The Convolve module







Rate Theory and CRESCENDO Sub module



- solute species can only be mobile if associated to a single defect \rightarrow (1,1) and (-1,1)
- · consistently, only single defect or single defect-solute pairs can be emitted



Rate Theory and CRESCENDO Sub module



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 \rightarrow some parameters related to this transformation and to the resolution numerical scheme can ne changed by (very) expert users



Tuning Rate Theory effective parameters



Tuning Rate Theory effective parameters





Obtain Critically Resolved Shear Stress (CRSS) from microstructural data

Orowan Bacon model for impenetrable obstacles accounting for dipole interaction between bending dislocation lines.

$$\tau_{\rm OB} = \left(\frac{\ln D'/b}{\ln l/b}\right)^{3/2} \frac{\mu b}{l} \frac{\ln(l/b)}{2\pi},$$

D: diameter of obstacle. C: concentration of obstacle. I is defined as $1/\sqrt{(DC)}$

$$\tau_{total} = \sqrt{\sum \tau_{OB,i}^2}$$

i running over all interstitial and vacancy clusters



G. Adjanor, et al., Journal of Nuclear Materials, 406, 175 (2010)

- all clusters are Orowan precipitates (with a self and mutual dipole interaction correction from the Bacon et al. model). Note that this assumption may lead to an overestimation of the pinning force in the case of copper precipitates and interstitial clusters and to an underestimation in the case of voids which seem to constitute the major source of hardening in typical irradiation conditions,
- the diameter of the clusters *D* is taken as spherical in all cases. This is actually the case for cavities, precipitates and heterogeneous precipitates, which are known to form atmospheres around cavities, but it is not the case for interstitial loops. Nevertheless, this approximation may not be too severe as their number density is small compared to that of cavities, and, again, because loops are also much weaker obstacles than cavities.

G. Adjanor, et al., Journal of Nuclear Materials,406, 175 (2010)

